

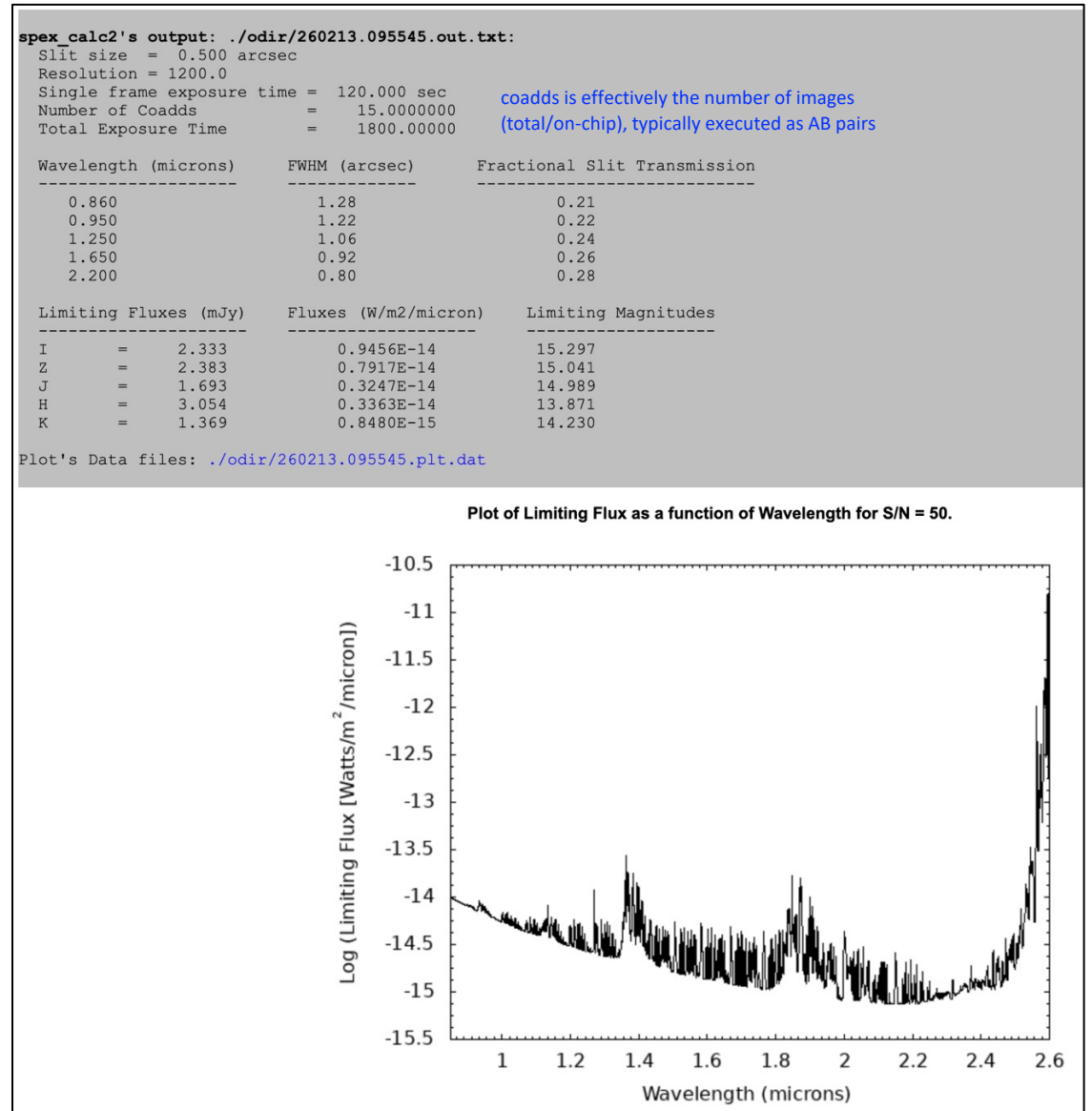
# SpeX

## 1. Limited Flux needs

- Observing mode
- Slit width
- Seeing
- S/N
- On-chip itime
- Total itime

Quantity to be estimated:	<input type="text" value="Limiting Flux"/>	
Select Observing Mode:	<input type="text" value="Short-Wavelength (0.8-2.5 microns) XD"/>	
Slit width (arcsec):	<input type="text" value="0.5"/>	
Seeing (FWHM) in arcsec at 2.2 microns:	<input type="text" value="0.8"/>	<u>Required for all</u>
Required Signal-to-Noise ratio:	<input type="text" value="50"/>	<u>Required for Limiting Flux, Total ITIME</u>
Single frame integration time (sec):	<input type="text" value="120"/>	<u>Required for all</u>
Total integration time(sec):	<input type="text" value="1800"/>	<u>Required for Limiting Flux, Signal-to-Noise</u>
Source Flux at 2.2 microns:	<input type="text"/> <input type="text" value="W/m2/micron"/>	Required for Total ITIME, Signal-to-Noise
Source blackbody temperature (K):	<input type="text"/>	Required for Total ITIME, Signal-to-Noise
<input type="button" value="Submit"/>	<input type="button" value="Reset"/>	

Simple to estimate limiting magnitude at any wavelength



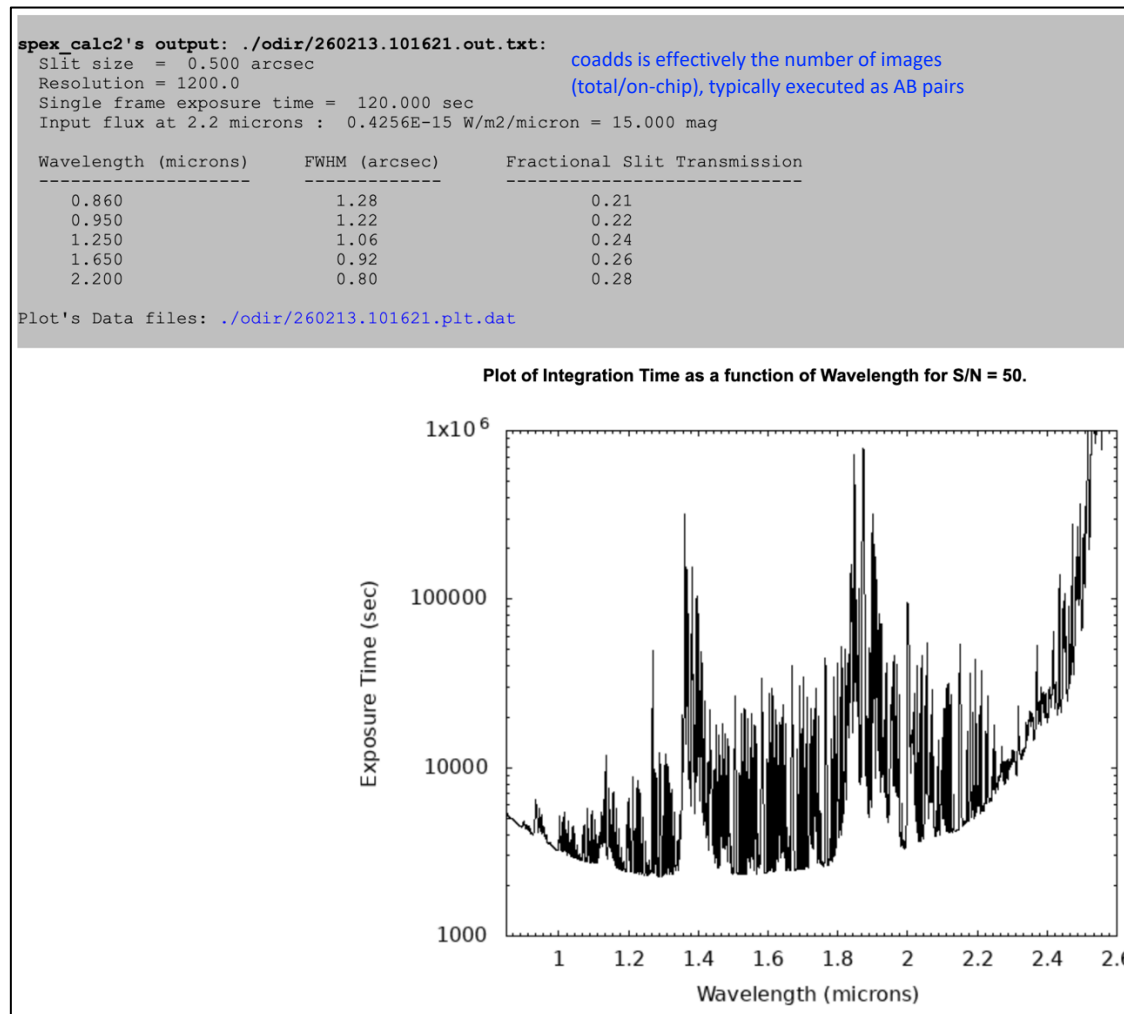
# SpeX

## 2a. Total Integration Time needs

- Observing mode
- Slit width
- Seeing
- S/N
- On-chip itime
- Source flux at 2.2  $\mu\text{m}$
- Estimated source BB temp

Quantity to be estimated:	<input type="text" value="Total Integration Time"/>	
Select Observing Mode:	<input type="text" value="Short-Wavelength (0.8-2.5 microns) XD"/>	
Slit width (arcsec):	<input type="text" value="0.5"/>	
Seeing (FWHM) in arcsec at 2.2 microns:	<input type="text" value="0.8"/>	<u>Required for all</u>
Required Signal-to-Noise ratio:	<input type="text" value="50"/>	<u>Required for Limiting Flux, Total ITIME</u>
Single frame integration time (sec):	<input type="text" value="120"/>	<u>Required for all</u>
Total integration time(sec):	<input type="text"/>	<u>Required for Limiting Flux, Signal-to-Noise</u>
Source Flux at 2.2 microns:	<input type="text" value="15.0"/> <input type="text" value="mag"/>	<u>Required for Total ITIME, Signal-to-Noise</u>
Source blackbody temperature (K):	<input type="text" value="5500"/>	<u>Required for Total ITIME, Signal-to-Noise</u>
<input type="button" value="Submit"/>	<input type="button" value="Reset"/>	

Useful to estimate the sensitivity simultaneously across a range of wavelengths given a particular source. In this example a 5500K BB source (G2V) is used



# SpeX

## 2b. Total Integration Time needs

- Observing mode
- Slit width
- Seeing
- S/N
- On-chip itime
- Source flux at 2.2  $\mu\text{m}$
- Estimated source BB temp

Quantity to be estimated:	<input type="text" value="Total Integration Time"/>	
Select Observing Mode:	<input type="text" value="Short-Wavelength (0.8-2.5 microns) XD"/>	
Slit width (arcsec):	<input type="text" value="0.3"/>	
Seeing (FWHM) in arcsec at 2.2 microns:	<input type="text" value="0.8"/>	<u>Required for all</u>
Required Signal-to-Noise ratio:	<input type="text" value="50"/>	<u>Required for Limiting Flux, Total ITIME</u>
Single frame integration time (sec):	<input type="text" value="120"/>	<u>Required for all</u>
Total integration time(sec):	<input type="text"/>	<u>Required for Limiting Flux, Signal-to-Noise</u>
Source Flux at 2.2 microns:	<input type="text" value="15.0"/> <input type="text" value="mag"/>	<u>Required for Total ITIME, Signal-to-Noise</u>
Source blackbody temperature (K):	<input type="text" value="1600"/>	<u>Required for Total ITIME, Signal-to-Noise</u>
<input type="button" value="Submit"/>	<input type="button" value="Reset"/>	

spex\_calc2's output: ./odir/260213.102042.out.txt:

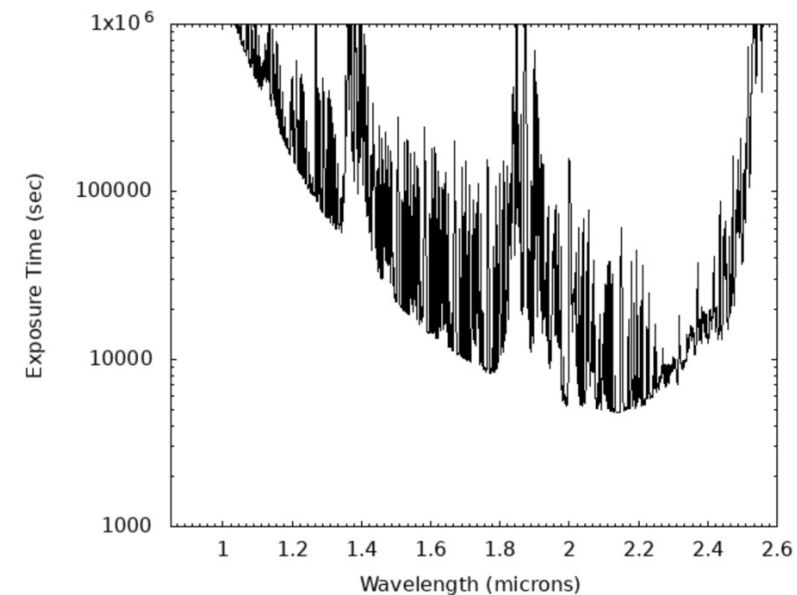
```
Slit size = 0.500 arcsec
Resolution = 1200.0
Single frame exposure time = 120.000 sec
Input flux at 2.2 microns : 0.4256E-15 W/m2/micron = 15.000 mag
```

coadds is effectively the number of images  
(total/on-chip), typically executed as AB pairs

Wavelength (microns)	FWHM (arcsec)	Fractional Slit Transmission
0.860	1.28	0.21
0.950	1.22	0.22
1.250	1.06	0.24
1.650	0.92	0.26
2.200	0.80	0.28

Plot's Data files: ./odir/260213.102042.plt.dat

Plot of Integration Time as a function of Wavelength for S/N = 50.



In this example a 1600K BB source is used. This also roughly simulates a visual extinction of  $A_v=10$  mags assuming an intrinsic BB temp of 5500K. (A 1000K BB also simulates a visual extinction of  $A_v=20$  mags assuming an intrinsic BB temp of 5500K). You will need to estimate the color temp corresponding to a desired  $A_v$  separately

# SpeX

## 3a. Signal-to-Noise needs

- Observing mode
- Slit width
- Seeing
- On-chip itime
- Total itime
- Source flux at 2.2  $\mu\text{m}$
- Estimated source BB temp

Quantity to be estimated:	<input type="text" value="Total Integration Time"/>	
Select Observing Mode:	<input type="text" value="Short-Wavelength (0.8-2.5 microns) XD"/>	
Slit width (arcsec):	<input type="text" value="0.3"/>	
Seeing (FWHM) in arcsec at 2.2 microns:	<input type="text" value="0.8"/>	<u>Required for all</u>
Required Signal-to-Noise ratio:	<input type="text" value="50"/>	Required for Limiting Flux, Total ITIME
Single frame integration time (sec):	<input type="text" value="120"/>	<u>Required for all</u>
Total integration time(sec):	<input type="text"/>	Required for Limiting Flux, <u>Signal-to-Noise</u>
Source Flux at 2.2 microns:	<input type="text" value="15.0"/> <input type="text" value="mag"/>	Required for Total ITIME, <u>Signal-to-Noise</u>
Source blackbody temperature (K):	<input type="text" value="1600"/>	Required for Total ITIME, <u>Signal-to-Noise</u>
<input type="button" value="Submit"/>	<input type="button" value="Reset"/>	

Useful to estimate the sensitivity simultaneously across a range of wavelengths given a particular source. In this example a 5500K BB source (G2V) is used

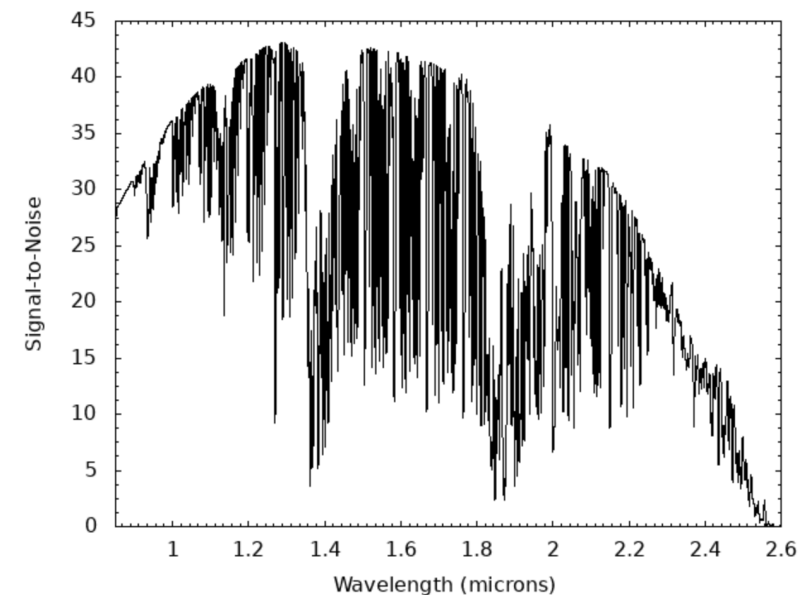
```
spex_calc2's output: ./odir/260213.010852.out.txt:
Slit size = 0.500 arcsec
Resolution = 1200.0
Single frame exposure time = 120.000 sec
Number of Coadds = 14.0000000
Total Exposure Time = 1680.00000 sec
Input flux at 2.2 microns : 0.4256E-15 W/m2/micron = 15.000 mag

coadds is effectively the number of images
(total/on-chip), typically executed as AB pairs
```

Wavelength (microns)	FWHM (arcsec)	Fractional Slit Transmission
0.860	1.28	0.21
0.950	1.22	0.22
1.250	1.06	0.24
1.650	0.92	0.26
2.200	0.80	0.28

Plot's Data files: ./odir/260213.010852.plt.dat

Plot of Signal-to-Noise as a function of Wavelength for S/N = .



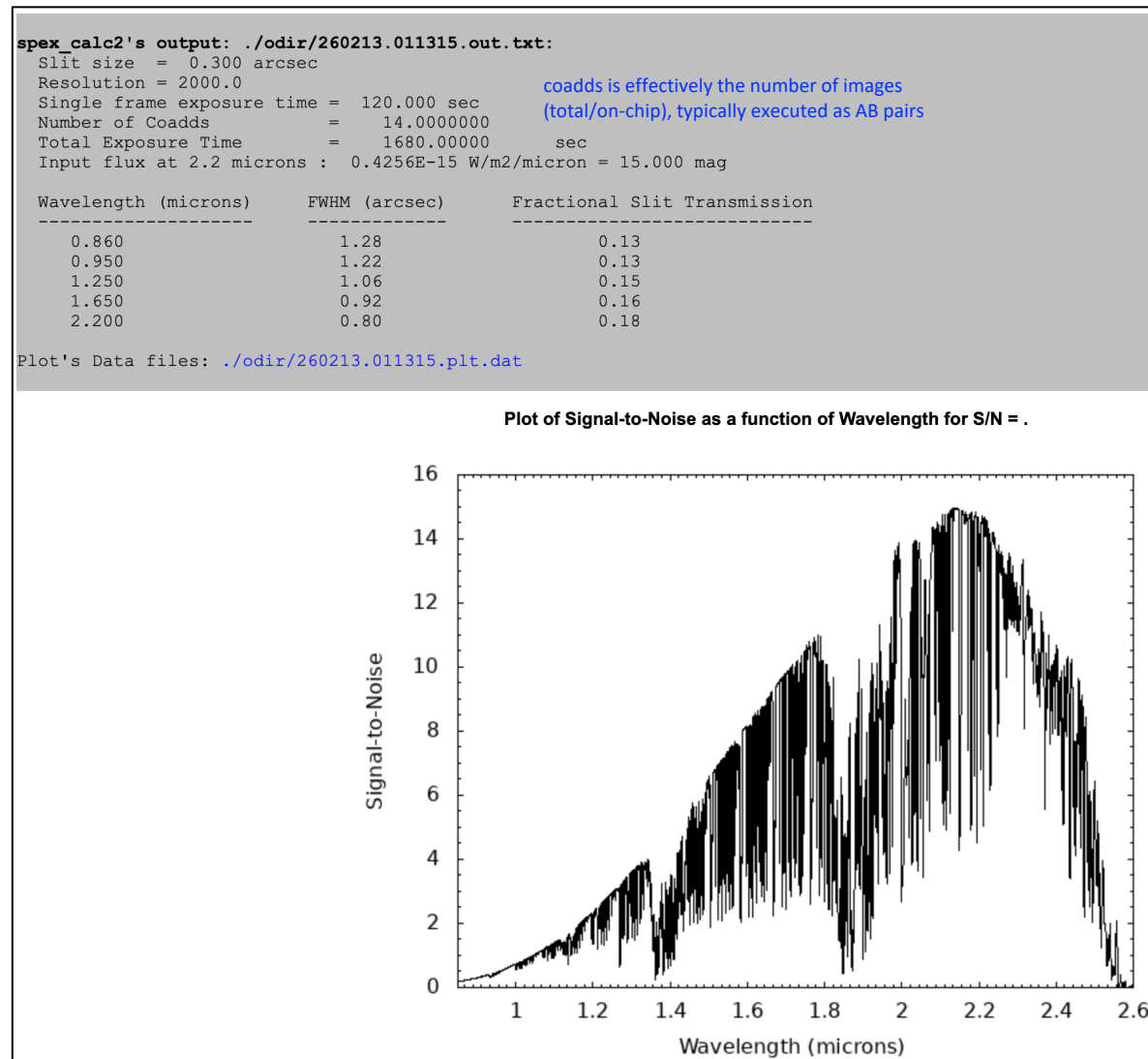
# SpeX

## 3b. Signal-to-Noise needs

- Observing mode
- Slit width
- Seeing
- On-chip itime
- Total itime
- Source flux at 2.2  $\mu\text{m}$
- Estimated source BB temp

Quantity to be estimated:	<input type="text" value="Signal-to-Noise"/>	
Select Observing Mode:	<input type="text" value="Short-Wavelength (0.8-2.5 microns) XD"/>	
Slit width (arcsec):	<input type="text" value="0.5"/>	
Seeing (FWHM) in arcsec at 2.2 microns:	<input type="text" value="0.8"/>	<u>Required for all</u>
Required Signal-to-Noise ratio:	<input type="text"/>	Required for Limiting Flux, Total ITIME
Single frame integration time (sec):	<input type="text" value="120"/>	<u>Required for all</u>
Total integration time(sec):	<input type="text" value="1680"/>	Required for Limiting Flux, <u>Signal-to-Noise</u>
Source Flux at 2.2 microns:	<input type="text" value="15.0"/> <input type="text" value="mag"/>	Required for Total ITIME, <u>Signal-to-Noise</u>
Source blackbody temperature (K):	<input type="text" value="1600"/>	Required for Total ITIME, <u>Signal-to-Noise</u>
<input type="button" value="Submit"/>	<input type="button" value="Reset"/>	

In this example a 1600K BB source is used. This also roughly simulates a visual extinction of  $A_v=10$  mags assuming an intrinsic BB temp of 5500K. (A 1000K BB also simulates a visual extinction of  $A_v=20$  mags assuming an intrinsic BB temp of 5500K). You will need to estimate the color temp corresponding to a desired  $A_v$  separately



# iSHELL

## iSHELL ETC needs

- Slit width
- On-chip itime
- Coadds
- Cycles
- Seeing
- SN

Slit Width (arcsec):

itime (sec)

coadds

cycles

seeing (arcsec)

SN

The iSHELL ETC is simpler than the SpeX ETC since iSHELL covers a much smaller one-shot wavelength range. Note that the ETC  $R$  for a given slit-width is slightly smaller than the actual values (see the iSHELL PASP paper)

## INPUT PARAMETERS

- Slit width (arcsec)= 0.750000
- SN= 100.000
- on-chip ITIME (sec)= 180.000
- number of coadds= 1.00000
- number of cycles= 10.0000
- $R$  (per spectral resolution element)= 38000
- Seeing FWHM at K (arcsec)= 0.800000

(1)	(2)	(3)	(4)	(5)
1.10um	9.4879374	9.9693583	2.3195521e-14	1.4887980e-14
1.25um	9.6790971	10.130847	1.4632401e-14	9.6519624e-15
1.65um	10.230420	10.623293	4.0385956e-15	2.8124260e-15
2.22um	10.022064	10.362918	2.3699773e-15	1.7314197e-15
2.50um	9.8260599	10.150282	2.3166142e-15	1.7185609e-15
3.00um	9.2049662	9.5097861	1.9703003e-15	1.4880052e-15
3.55um	8.3970556	8.6922990	2.7355749e-15	2.0842570e-15
4.00um	7.0776486	7.3722780	6.2126507e-15	4.7361470e-15
4.77um	5.7064570	6.0121534	1.3817658e-14	1.0426916e-14

- (1) Wavelength (microns)
- (2) Point source magnitude (Vega)
- (3) Extended source magnitude per square arcsecond (Vega)
- (4) Point source line flux (erg per second per square cm)
- (5) Extended source line flux (erg per second per square cm per arcsec)